In an infinite medium, the total field $\psi(\mathbf{r})$ (velocity potential, or pressure if you want) is given by the volume integral

$$\psi(\mathbf{r}) = \int_{V_0} f(\mathbf{r}_0) g_\omega(\mathbf{r} | \mathbf{r}_0) \,\mathrm{d}v_0 \tag{1}$$

where $f(\mathbf{r}_0)$ is the source strength, and where $g_{\omega}(\mathbf{r}|\mathbf{r}_0)$ is the free-field Green's function.

Noting that the total field is also given by the Helmholtz-Kirchoff integral,

$$\psi(\mathbf{r}) = \iiint f_{\omega}(\mathbf{r}_0)g_{\omega}(\mathbf{r}_0|\mathbf{r})dv_0 + \oiint \left\{g_{\omega}(\mathbf{r}_0|\mathbf{r})\frac{\partial\psi(\mathbf{r}_0)}{\partial n_0} - \psi(\mathbf{r}_0)\frac{\partial g_{\omega}(\mathbf{r}_0|\mathbf{r})}{\partial n_0}\right\}dS_0,$$
(2)

Comparing equations (1) and (2) shows that the surface integral in (2) must vanish:

The area of a sphere goes as R^2 , so for the right-hand-side to go to 0, the integrand must go to 0 faster than R^{-2} as $R \to \infty$. Since $\frac{e^{ikR}}{4\pi}$ is oscillatory, it does not affect the convergence to 0. The condition for convergence is therefore¹

$$\lim_{R \to \infty} \left(\frac{1}{R}\right) \left(\frac{\partial \psi(\boldsymbol{r}_0)}{\partial n_0} - ik\psi(\boldsymbol{r}_0)\right) \div R^{-2} = 0$$
$$\lim_{R \to \infty} R\left(\frac{\partial \psi(\boldsymbol{r}_0)}{\partial n_0} - ik\psi(\boldsymbol{r}_0)\right) = 0$$
$$\lim_{R \to \infty} R\left(\frac{\partial}{\partial n_0} - ik\right)\psi(\boldsymbol{r}_0) = 0$$

This is the so-called Sommerfeld radiation condition, which must be satisfied by any field in an infinite medium.

¹For two functions A(x) and B(x) that approach 0 as $x \to \infty$, $A(x) \to 0$ faster than B(x) if $\lim_{x\to\infty} \frac{A(x)}{B(x)} \to 0$.